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W!!!em T. Slajton<br>Bitcrowave Antennes ard Components Branch Electronles D!vision

November 0, 1954


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#### Abstract

ARSTRACT A set of antenna gain-standard horns covering the milcrowave range from 0.77 cm to 31.5 cm has been designed aud carefully call-  be dupifcated accu:ately from the drawings supplied. A stople method of extending and improving the aecuracy of Scheicunoft's gatn curves is also described.


PROBLEM STATUS
This is a final report on this piaase of the problem; work on the problem is coctinuing.

## AUTHORIZATION

NRL Problem R09-03
Project NiR 888-03C

Manuseript submitt:d September 2, 1954

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## DESIGN AND CALIBRATION

## OP ? Mitiowaye antenna gain standards

## INTRODUCTION

Tha need fox accurate and practlcal microwave antenna gain standards has led to the design and calibration of a series of pyramidal horns covering the microwave bands
 Eanging from 24.7 db to 13.7 db . There is 2 horn for each wavegulde size in the rante. The l:orns can be easily and accurately duplleated from cirawings supplied in tils report.

## DESign

Titree requiremants were considered of prime importance in the design: a useful gain fizare, simplicity of construction, and accuracy of callbration. The fabricated type of hora (rig. 1), with flat metal shppis forming the sides, was resided upon as the best means of satisfying the construction reguirements. For simplicity, the horns werdesigned so that the E - and H -plane flares meet the wivegulde is a common plane.

Another consideration was the over-all size and welgit. It was impractical to scale the horns from ona band to another ihroughout the range, since the horns at the loriger wavelengths would be too large and those at the shurte= wavelengths too small. Accordingly, there are five difierent designs; each of the other six horns was' scaled from one of these.

The $8-\mathrm{mm}$ and $1.8-\mathrm{cm}$ horns were scaled from the $1.25-\mathrm{cm}$ horn; the $4.75-\mathrm{cm}$ horn from the $3.2-\mathrm{cm}$ horn; the $3.95-\mathrm{cm}$ and $6-\mathrm{cm}$ horns from the $10-\mathrm{cm}$ horn; and the $15-\mathrm{cm}$ horn fiom the $23-\mathrm{cm}$ horn. In scaling, the values of $\ell_{\mathrm{H}}$ had to bealtered slightly in order to make a sicmple junction at the wavegutde. This was necessary because, Whith one or two exceptiens, the inside dimensions of the waveguldes ave not ecaled from ne band to another. Tho 2c.justment made only a very sllght changs in. the calcuiated gatin (aiourt 0.02 to 0.03 db ).


Fig. 1 - Physical dimensions, for calculating the ga'n

The $3.95-\mathrm{em}$ horn reprosents an ceverlapning of the $3.2-\mathrm{cm}$ band and the $4.75-\mathrm{cox}$ band. Eorruyer, it was deeided to taclude tris horn in the serles because it fits a standard
 chaciss on the $10-\mathrm{cm}$ horn frow which it is scaled.

The basic design data Inclucing the dimensions, operating range, and tesign-point gain for all the horns are summarized In Table A-2.*

Readers who are interested in a detailed design procedure are 5 -uferred to the Appendix, where a simple means of extending the range of Schelkunoft's gain curves and Improving the accuracy of the gain flgure ottainable from them ts described. This methed eliminates the necessity for long compatations involving Fresnel Integrals, and yields very close agreement with the detailed calculations.

## CONSARUCTION

As mentioned previously, the fabricated type of horn using flat metal si :ets was decided upon as most suitajle. The one exception is the 8 -mm design, where electroforming was considered necessary because of the small size and close tolerances. Horns for the hands from 1.25 cm to 10 cill were mad= of bases sheets. At the $15-, 23-$, and 3 n-rm bands, hornc wore fabringten from sheet aluminum using hellum gas to facllitate welding the Joints (heliarc process). This construction reduced üe weigit soaidicrably and was icund to be satisfactory for proulucing accurate, vaiform, and augged horns.

Dimensiens for each set of horns are given in Figs. A-6 through A-17.

## CAlibration

Experimental primary gain measurements (Fig. 2) were maje In order to check the accuracy of the calculated gain.t rreat care was taken in maklou fiese measuremenis. Both the horns and the bolometer detectors were carefully matched and the bolometer amplificr and guigut seier (VTVM) were calibrated accurately. The bolometer amplifier was found to be tinear throughout the range used. The use of $r-f$ coaxial cables was avolied because of instaillity, wavegulde being used inctead. Microwave absorbent material (1) nas used to minimize reflections. Even so, duficulties vere encountered at the longer wavelengths because of reflections and the large separation distances required. As Braun has shown (2), true Fraunhoter fleld conditions do not exist untll a separation distance between horns of many times $2 \mathrm{~d}^{2} / \lambda$ is attained, d being the larger aperture dimension. Because of these dificuli:eg, axpeimential galn measurements at 10 cm and above wore abandoned. It was decided to scale the $3.95-\mathrm{cm}=8 \mathrm{~d} 8 \mathrm{~cm}$ herns from the $10-\mathrm{cm}$ horn in order to obtain rellable measurements at the shorter wavelengths. Figure 3 shows the anechole test site. An example of the method used in ovaluating the experimental data is given in the Appendix.


Fig. 2 - Experimental setup for gain measurements

[^0]

Fig. 3 - Anechoic test site

Measurtiduts were mafe at several scparation distances in each case, and were repeated many times, changians such rar'zhles as the power level and the peaking of the hurns. See figs. 3 and 4.

$r:=$ - Horn and tearsmitter on adjuztable mount $\therefore \therefore$ :
cain curves for eaen band are shown in Flg. A-5 (a,b, c). Figares A-4 (a-f) ehotw the field patterns for three basic horn designs.

## REMARKS

Horms representing icur iasic designs were measured for mismatch over their bands. The greatest VSWR's enccuntered in the varlous bands are as follows:

| Band | Max VSWR |
| :---: | :---: |
| 1.8 cm | 1.10 |
| 3.2 cm | 1.20 |
| 6 cm | 1.25 |
| 23 cm | 1.20 |

The horns for the other bands should have a VSWR ciose to that of the horns irom which they were scaled.

In any event, when the horns are used in gain measurements, the VSWR shnuld he ㅍëēsurei ui une waveiength used, and for accurate measurements the horns ahould be =arefully matched, or aliowance shouid be made for any teismatch. In either case the bolometer must be well-matched. The use of flange-to-flange connections rather than chokes, is recommended whenever operating at a wavelength dufering from that for Which the shokes were designed, since at some wavelengthe choke-to-flange joln's may introduce considerable m!smatch.

## Accuracy

At any one wavelength the measured polnts showed a dispersion of less than 0.1 db. As a function of wavelength, the gain curve is not monotonic, as qould be predicted from the theory, but shows small, though definite, periodic wiegles (see Fig. A-5 (b)j. After exhaustive checking it is ife!t that these wiggies are actually present, and not due to experimental difilculties. Thiss efiact can probably be attributed to higher modes in the aperture and curients on the outside of the horn, both of risish are sigiected In the theory. However, since the wiggles are small, and since a tremendess amount of additional data would have to be taken to reprodicen the wiggles accurately, a curve drawn through the average of the measured yoints was uscd. Taking Into account all posslive deviat!ons from the true gain over each band, it was decided that the marimum possiblo error would te less thun $\pm 0.3$ th up to and including $:^{3}-710$-cm horns.

At wavelengths longer than 10 cm , where no direct experimental checks have been feasible, the gain has been calculated by means of Schelkunot!'s formula. To arrivo at u reasonatle tolerance at these iavelength, it was note: that below 10 cra the groatest
 near fleid elfects) and the calcilated gain at the same wavelength was of that order of 0.2 db . In general the difference wis unch less than dis lifure. Therefore it is folt that a tolorance of $\pm 0.5 \mathrm{db}$ is reasonable for all horns above the $10-\mathrm{cm}$ band. In all probahillty, the actual errors are consicierably less than the maximum possible toleraness quoted.

## ACKNOWLEDGMENTS

The author vishes to express his apprecfation to E. H. Braun for his advice and cooperation and to F. W. Lashuay for his suggestions in connectlon with the cunstruction of the horns.

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## LIST CF APPENDDX ILLUETRATIC.!3

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## APPENDIX

Methods for Determintig Horn Dimeasions and Galn

## BACKGROUND

Schelkunoif's gain curves in various forms $1,2,3$ were used for determining the tentative dimensions of the horns and for obtaining a first approximation to the galn. After the aperture dinensions had been chosen and a zeasonable value for $\ell_{E}$ (the $E$-piane slant height) had been set, the $\mathcal{F}$-plane slant helght, $\mathcal{I}_{\mathrm{H}}$, was uniquely deterinined by the requirement that the flared sides of the horn meet the waveguife in the same plane (Fig. 1, p. 1). For the purpose of calculating tie expected gain, this value of $\ell_{H}$ was approximated by the relation:
where $=$ H-plane aperture dimession
$\mathbf{b}=\mathbf{E}$-plane aperture citmension
$\nabla_{c}=$ E-plane taside ilmension of the wavegulde
$\sigma_{B}=H$-plane instue dimension of the maveguide.
After the tentative gain had ceen determined, the exact value of $\ell_{h}$ was obtalned from the formula

$$
\begin{equation*}
t_{H}=\frac{a}{-v_{2}} \sqrt{\left[\left(L_{c}\right)^{2}-\left(\frac{b}{2}\right)^{2}\right]\left[\left(1-\frac{v_{s}}{b}\right)^{2}\right] \div\left[\frac{-m_{n}}{2}\right]^{2}} \tag{2}
\end{equation*}
$$

[^1]In uslag Schelrunoif's gain curves, it was found that no one family of curves in the zefer .es nuentloncd covered 2 range great enough to include all the desircd sizes of horno. Errthermore, certain parts of the curves were found to be less accurate than othe:s. To overcome these difficu:ties a new procedure has been devised. 4 A briel revict - . relationship of the curves to the gain formula will help to clarify the procedurt. in ar notation is substant'plly that used in the recent book by Schelkunoff and Frils, ${ }^{3}$ a: - : y Sllver.2

The Schelkunoff curves give the directive gain for horns flared in either of the two principai piarcs; $g_{g}$ is ti:e directive gain of a sectoral horn flared in the $\overline{\mathrm{E}}$-plane, and $\mathrm{g}_{\mathrm{n}}$ is the dircctive gain of a sectoral norn ilared in the H-pianc. The two secto:al gain curves are obtained irom the following formulas, expressed in terms of the tabulated Fresnel integrals $C(X)$ and $S(X)$ :

$$
\begin{align*}
& \frac{\lambda}{\delta} g_{H}=\frac{4 \pi l_{H}}{v}\left[\{c(u)-c(v)\}^{2}+\{s(u)-s(v)\}^{2}\right]  \tag{3}\\
& \frac{\lambda}{d} z_{z}=\frac{s 4 l_{2}}{\pi}\left[c^{2}(v)+s^{2}(v)\right], \tag{4}
\end{align*}
$$

where

$$
\begin{aligned}
& u=\frac{1}{\sqrt{2}}\left(\frac{\sqrt{\lambda \ell_{G}}}{a}+\frac{a}{\sqrt{\lambda \ell_{B}}}\right) \\
& v=\frac{1}{\sqrt{2}}\left(\frac{\sqrt{\lambda \ell_{R}}}{a}-\frac{a}{\sqrt{\lambda \ell_{B}}}\right) \\
& z=\frac{b}{\sqrt{2 \lambda \ell}} \\
& \lambda=\text { wavelsngth. }
\end{aligned}
$$

The gain of a pyramidai horn is

$$
E=\frac{8 \pi \ell_{E} \ell_{i 1}}{b b}\left[c^{2}(w)+r^{2}(w)\right]\left[\{c(u)-\alpha(v)\}^{2}+\{S(u)-S(v)\}^{2}\right]
$$

This resuit can easily be obtalned from the tro sectoral curves by multiplying togethor
 formula

$$
\begin{equation*}
g=\frac{\left(\frac{\lambda}{4} \varepsilon_{E}\right)\left(\frac{\lambda}{h} \varepsilon_{1 H}\right)}{\frac{32}{-\pi}} \tag{5}
\end{equation*}
$$

where $\frac{\lambda}{a} x_{z}$ and $\frac{\lambda}{b} g_{n}$ are read ulrectly from the curves.

[^2]
## FXTENSION AND APPLICATION

Braun's methof ${ }^{4}$ provides a conyenlent means of extending the range of the gain curves and eliminating the Inaccuracy arlaing from interpolatlons between curves. He intruduces the arbitrary factors $k_{E}$ and $k_{H}$ to create a fictitious horn having these ditmensions:

$$
\begin{array}{ll}
t=k_{H} A, & \ell_{H}=k_{H}^{2} L_{H} \\
b=k_{E} B, & \ell_{E}=k_{E}^{2} L_{E}
\end{array}
$$

where $A, B, L_{E}$, and $I_{H}$ are the actual horn dimar.sions. By choosing the proper value for $k_{z}$ and $k_{k}$, one can make $l_{E}$ and $\ell_{H}$ fall exactly on one of the respective galn curves for each plane. Aftor the galin of the 'Ietitous hern lifict, 1 is read from the curves, the gain of the actual horn (ract.) is obtained from the relation

$$
g_{\text {act }}=\frac{E_{\text {fict }}}{k_{z} k_{R}}
$$

 The Schelkunolf corves for $\ell_{z}=50 \lambda$ and $\mathcal{L}_{4}=50 \lambda$ are convenient for this purpose and have been accurately recomputed and plotted on an expanded scale in Figs. A-2 (a,b) and A-3 (a,b) so that they may be read with such accuracy that it is no longer necessary to make the dciatled calculat!ons invelyed in using the ailn formula. The curves wero ploted from formulas (3) and (4). The values obtatned frcm these formulas are tabulated in rabia i-1. For maxiaum accuracy these values may be preferable to those obtatnod from the curpes. Linear Interpolation between points will yicld good ircuiacy. The table makes it possible to plot any desired portions of the curves on whatever scale is preferred.

An example will demonstrato the simpliclty of the mothod.

$$
\begin{array}{lll}
\text { Actual horn: } & A=8.13 \lambda, & L_{M}=19.72 \lambda \\
& B=6.67 \lambda_{\mathrm{E}} & \mathrm{~L}_{\mathrm{E}}=18.52 \lambda
\end{array}
$$

If it desired to maks use of the $50-\lambda$ curves referred to above, the $k$ ' are chosen as follows:

$$
\begin{aligned}
& k_{2}^{2}=\frac{50 \lambda}{i 8.52 \lambda}=2.6998, \quad k_{2}=1.643, \\
& k_{H}^{2}=\frac{50 \lambda}{10.72 \lambda}=2.5355, \quad k_{H}=1.592 .
\end{aligned}
$$

Fictitious horn: $b=k_{z} B: 10.96 \lambda_{1} \ell_{z}=50 \lambda_{4}$

$$
==k_{n} A=12.54 \lambda . \ell_{H}=50 \lambda .
$$

From the so- $\lambda$ galn curvas

$$
\begin{aligned}
& \frac{\lambda}{k} R_{R}=80.77 \\
& \frac{\lambda}{b} R_{n}=98.92 .
\end{aligned}
$$

From formula (5),

$$
\begin{aligned}
& \mathrm{R}_{\text {fict. }}=\frac{\left(\frac{\lambda}{\left.\frac{\lambda}{R_{Z}}\right)}\left(\frac{\lambda}{b} \kappa_{H}\right)\right.}{\frac{32}{\pi}}=784.40 . \\
& \mathrm{g}_{\text {act. }}=\frac{\mathrm{g}_{\text {fict. }}}{k_{Z} k_{H}}=299.38, \text { or } 24.77 \mathrm{db} .
\end{aligned}
$$

Detalled calculations using the Fresnel integrais in the cain formula resulted in the game gain figure, 24.77 db . Similar comparisons at each of the other bands showed agreenent within 0.01 db .

## USE OF CORRECTION CURVES

The procedure for determining the true Fraunhoier gain from the primary gain test data, using Braun's near field correction curves, Fig. A-1 ( $a, b$ ), is shown in the following example taken Irom acius! measuremenis:

X-band horn cimenslons: $\quad=7.654 \mathrm{in} ., \quad \ell_{n}=13.484 \mathrm{in}$.

$$
b=5.669 \mathrm{in} ., \quad \ell_{x}=12.598 \mathrm{in} .
$$

$\lambda=3.20 \mathrm{~cm}=1.2598 \mathrm{in}$.
R Idistance between horn=: = 140.25 in.
$\frac{4 \pi \mathrm{R}}{\lambda}=\frac{i 12.566)(140.25)}{1.2539}=1398.9$.
Frome test data $\frac{P_{T}}{P_{R}}=\frac{11.3}{0.523}=91.87: \quad \sqrt{\frac{P_{T}}{P_{R}}}=9.585$
where $P_{T}$ represents power transmitted and $P_{R}$ power recelved.
$\operatorname{cosin}_{\text {uncorrected }}=\frac{\frac{4 \pi R}{\lambda}}{\sqrt{\frac{P_{I}}{P_{R}}}}=\frac{1398.9}{9.585}=245.95$, or 21.54 dh .
Parameters for using the correction curves:
E-plane:

$$
\left.\begin{array}{l}
\frac{8 l_{z}}{b^{2}}=\frac{(8)(12.598)}{32.17^{2}}=3.1360 \\
E=\left(\frac{9 l_{x}}{b^{2}}\right) \quad \lambda i=(3.1360)(1.2598)=3.951 \\
\log \frac{2 k}{b^{2}}=\log \frac{(1.2598)(140.25)}{32.13}=\log 5.498=0.740
\end{array}\right\}
$$



Fig. A-1 - Brann's E- and Il-plane correction curves

H-plane:

$$
\left.\begin{array}{rl}
\frac{8 \ell_{H}}{R^{2}} & =\frac{(3)(13.434)}{58.584}=1.8413 \\
H & =\left(\frac{3 L_{n}}{a^{2}}\right) \lambda=(1.8413)(1.2598)=2.320 \\
\log \frac{\lambda R}{R^{2}} & =\log \frac{(1.2598)(140.25)}{58.584}=\operatorname{loR} 3.016=0.479
\end{array}\right\}
$$

## Reading from the correction curves:

$$
\text { E-piàné curreciion ................................................................................................. } 0.22 \text { db }
$$

H-plans correction ................................................................................................. 0.28 db

Uncurrected gain (above) ................................................................................................ 21 db


The calculated gain, using Schelkunoff's formula, in this case was the same: 22.14 db .

## DETERMINATION OF AN OPTLHUA: HCRN WITH <br> SPECIFIED GAIN AND EQUAL BEAMWDTHS

A. simule means has been devised for finding the dimensions of a horn which satisfles the following requirements:
(1) Spectlied gain
(2) Critmum horn*
(3) Equal heamwidths at the hall-power points.

Although this can be done mnipirlcally, $z$ sct of factors was determined . 50 m Schelkunof's gain formula, which yield the required horn parameters as a function of the absolute gain, g, alone.t These are as follows:

[^3]$\frac{a}{\lambda}=0.4675 \quad \sqrt{6}$
$\frac{b}{\lambda}=0.3463 \quad \sqrt{8}$
$\frac{f_{5}}{\lambda}=0.05764 \quad \mathrm{E}$
$\frac{\ell_{\pi}}{\lambda}=0.06885$
where $t, b, \ell_{\mathrm{E}}$, and $\ell_{\mathrm{H}}$ are t.z usual parameters as defined (p.7).
A horn raving these dinensions will have exactly the desired theoretical galn, and will be exactly an optimum hom. Howeyer, it should be pointed out that where a simpie joint between the flared horn azd the wavegulde is desired, the value of $\boldsymbol{d}_{\mathrm{H}}$ must le modtfled to make the horn fit the g:ide. Thls will change the gain by a small amount, usually a few tenths of a db, since the torn will no lorger be exaclly optimus., If a discrepancy
 formuia (2).

When a cioser approach to the specified gain is desired, a slight clange in the procedure is necessary. This is accomplished by the foliowing steps:
(1) Ccmpute tentative parameters $a^{\prime}, b^{\prime}$, and $\ell_{z}{ }^{\prime}$ in the same way as a, b, and $\ell_{E}$ were computed above.
(2) Obtain the approximate value, $\ell_{q}{ }^{\prime}$, to fit the wavegulde, using formula (1), p.7.
(3) Calculate the tentative gain, $x^{\prime}$, by the method outilined on p. 9 using

(4) Recompute $a, b$, and $\mathcal{L}_{2}$, substituting $\boldsymbol{\varepsilon}^{2} / g^{0}$ for $:$
(5) Obtaln the exact value of $\ell_{n}$ from formula (2
(6) Recaiculate the gain for the new pirameters.

Since the theorciscal gin is obiainen fory accurately in step 8 , it is easy to determine the discrepancy between the cesired gain and that now resulting from the adjustment to fit the wavegulde.


Fig. A-2 (a) Expanded E-plane theoretical gain curve


Fig. A-2 (b). Expanded E-plane theoretical gaincurve
$x_{2}$


Fig. A-3 (a). Expanded H-plans therretical gain curve


Fig. A-3 (b). Expanded If-plane theorctical gatn curre

TABIE A -1
Datin for Theoretical Gain Curves

| (a) E.Plane ( $L_{z}=50 \lambda$ ) |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $b$ | $\frac{\lambda}{4} \mathrm{~F}_{Z}$ | b |  | b | ${ }^{\lambda}{ }^{\text {a }} \mathrm{g}_{\mathrm{E}}$ | b | ${ }^{\boldsymbol{\lambda}} \mathbf{k}_{\mathbf{E}}$ | $b$ | 寅 8 E | $b$ | ${ }^{\lambda}{ }^{\text {E }}$ E | $b$ | $\frac{\lambda}{13} \mathbf{E}_{E}$ |
| 2. | 20.362 | 4.6 | 46.397 | 7.2 | 69.123 | 9.8 | 83.301 | 12.4 | 73.884 | 15.0 | 46. 493 | 7.6 | 19.910 |
| 2.1 | 21.381 | 4.7 | 47.362 | 7.3 | 69.847 | 9.9 | 81.426 | 12.5 | 13.041 | 15.1 | 45.268 | 17.7 | 19.316 |
| 2.2 | 22.395 | 4.8 | 48.326 | 7.4 | 70.555 | 10.c | 81.518 | 12.6 | 72.265 | 15.2 | 44.040 | 17.8 | 18.767 |
| 2.3 | 23.410 | 4.9 | 49.233 | 7.5 | 7: . 248 | 10.1 | 81.581 | 12.7 | 71.452 | 15.3 | 42.813 | 17.9 | 18.264 |
| 2.4 | 24.425 | 5.0 | 50.233 | 7.6 | 71.923 | 10.2 | 81.611 | 12.8 | 70.621 | 15.4 | 41.593 | 18.0 | 17.805 |
| 2.5 | 25.440 | 5.1 | 51.181 | 7.7 | 72.586 | 10.3 | 81.609 | 12.9 | 69.753 | 15.5 | 40.379 | 18.1 | 17.395 |
| 2.6 | 26.456 | 5.2 | 52.123 | 7.8 | 73.219 | 10.4 | 81.575 | 13.9 | 68.856 | 15.6 | 39.1.4 | 18.2 | 17.030 |
| 2.7 | 27.472 | 5.3 | 53.057 | 17.9 | 73.841 | 10.5 | 81.510 | 13.1 | 67.931 | 15.7 | 37.552 | 18.3 | 16.714 |
| 2.8 | 28.481 | 15. | 53.985 | [8. | 74.441 | 10.6 | 81.408 | 13.2 | 66.980 | 15.8 | 36.801 | 18.4 | 16.445 |
| 2.8 | 29.490 | 15. | 54.908 | 18 | 75.025 | 10.7 | 61.277 | 13.3 | 65.091 | 15.3 | 35.635 | 18.5 | 16.223 |
| 3.0 | 30.503 | 15.6 | 55.821 | S. 2 | 75.585 | 10.8 | R1. 110 | 13 | 04.997 | 16.0 | 34.488 | 18.6 | 16.048 |
|  | 31.511 | 5.? | 56.720 | 0.3 | 74.127 | $: 9.2$ | g0. 203 |  | 63.9E\% | 16.1 | 33.350 | $: 8.7$ | :5. $22:$ |
| 3.2 | 32.518 | 5 | 57.626 | 2.4 | 76.645 | 11.0 | 80.676 | 13.6 | 62.917 | 16.2 | :2.250 | 18.8 | 15.839 |
| 3.3 | 33.527 | 5.9 | 56.517 | 18.5 | 77.142 | 11.1 | 80.405 | 13.7 | 61.844 | 16.3 | 31.164 | 18.9 | 15.804 |
| 3.4 | 34.530 | 6.0 | 59.401 | 8.6 | 77.616 | 11.2 | 80.104 | 13.8 | 60.748 |  | 30.104 | 19.0 | 15.812 |
| 3.5 | 35.534 | 5.1 | 50.272 | 10. 7 | 75.065 | 11.3 | 79.765 | 13 | 59.635 | 16.5 | 29.069 | 19.1 | 15.870 |
| 3.6 | 36.534 | 6.2 | 61.134 | 8.8 | 78.492 | 11.4 | 79.393 | 14.0 | 58.501 | 16.6 | 28.063 | 19.2 | 15.967 |
| 3.7 | 37.531 | 6.3 | 61.987 | 8.9 | 78.892 | 11.5 | 78.987 | 14.1 | 57.351 | 16.7 | 27.086 | 19.3 | 16.108 |
| 3.8 | 38.530 | 6.4 | \|62.828 | 19.0 | 79.269 | 11.6 | 78.545 | 14.2 | 56.188 | 16.8 | 26.142 | 19.4 | 16.289 |
| 3.9 | 39.524 | 6.5 | 63.659 | 10. 1 | 79.619 | 11 | 78.068 | 14.3 | 55.008 | 16.9 | 25.232 | 19.5 | 16.521 |
| 4.0 | 40.515 | 6.6 | 64.477 | 9.2 | 79.944 | 11.8 | 77.559 | 14.4 | 53.816 | 17.0 | 24.355 | 19.6 | 16.769 |
| 4.1 | 41.504 | 6.7 | 65.285 | 9.3 | 80.240 | 111.9 | 77.014 | 14.5 | 52.614 | 17.1 | 23.515 | 19.7 | 17.064 |
| 4.2 | 42.490 | 6.8 | 66.089 | 9. | 80.510 |  | 76.435 | 14.6 | 51.402 | 17.2 | 22.713 | 19.8 | 17.394 |
|  | 43.472 | 6.9 | 66.862 | 9.5 | 80.752 |  | 75.822 |  | 50.183 | 17.3 | 21.951 | 19.9 | 17.755 |
|  | 44.450 | 7.0 | 67.630 | 9.6 | 80.964 | 12.2 | 75.176 74.897 |  | 78.959 | 17.4 | 21.228 | 20.0 | 18.147 |
|  | 45.425 | 7.1 | 68.385] | 9.7 | 81.146 | 12.3 | 74.497 |  |  | 17.5 | 20.548 |  |  |
| (b) H-Plane ( $\hat{H}_{H}=50 \lambda$ ) |  |  |  |  |  |  |  |  |  |  |  |  |  |
| - | ${ }_{\frac{1}{6}}^{\mathbf{E}_{2}}$ |  | ${ }^{\frac{1}{b}}{ }_{4}$ |  | $\frac{\lambda}{\lambda} \mathrm{E}_{1}$ | - | $\frac{1}{6} E_{n}$ |  | $\frac{1}{6} \mathrm{~g}_{4}$ |  | $\frac{1}{k}{ }^{\text {a }}$ |  | $\frac{1}{b} \mathrm{E}_{1}$ |
| 2. | 20.370 | 4.6 | 46.635 | 7.2 | 71.291 | 9.8 | 90.533 | 12.4 | 99.019 | 15.0 | 92.591 | 17.6 | 75.416 |
| 2.1 | 21.387 | 4.7 | 47.628 | 7.3 | 72.164 | 9.9 | 21.195 | 12.5 | 199.052 | 15.1 | 92.066 | 17.7 | 74.701 |
| 2.2 | 22.402 | 4.8 | 48.619 | 7.4 | 73.031 | 10.0 | 91.740 | 12.6 | :99.062 | 15.2 | 91.529 | 117.8 | 73.991 |
| 2.3 | 23.422 | 4.9 | 49.609 | 7.5 | 73.889 | 10.1 | 92.270 | 12.7 | 199.051 | 15.3 | 90.972 | 17.9 | 73.282 |
| 2.4 | 24.439 | 5.0 | 50.595 | 7.6 | 74.739 | 10.2 | 92.781 | 12.3 | 99.012 | 15.4 | 50.4c0 | 18.0 | 72.581 |
| 2.5 | 25.452 | 5.1 | 51.578 | 7.7 | 75.580 | 10.3 | 93.274 | 12.9 | 98.053 | 15.5 | 89.822 | 18.1 | 71.886 |
| 2.6 | 26.472 | 5.2 | 52.559 | 7.8 | 76.413 | 10.4 | 93.751 | 13.0 | 98.871 | 15.6 | 89.214 | 18.2 | 71.199 |
| 2.7 | 27.488 | 5.3 | 53.536 | 7.9 | 77.236 | 10.5 | 94.208 | 13.1 | 98.763 | 15.7 | 88.601 | 18.3 | 70.516 |
| 2.8 | 28.508 | 5.4 | 54.512 | 8.0 | 78.049 | 10.6 | 94.646 | 13.2 | 98.638 | i5.8 | 87.976 | 18.4 | 69.847 |
| 2.9 | 29.518 | 5.5 | 55.475 | 8.1 | 73.854 | $10 \overline{.} 7$ | 95.067 | 13.3 | 98.406 | 15.9 | 87.337 | 18.5 | 59.183 |
| 3.0 | 30.532 | 5.6 | 56.449 | 8.2 | 79.644 | 10.8 | 95.470 | 13. | 98.309 | 16.0 | 86.688 | 18.6 | 68.534 |
| 3.1 | 31.545 | 5.7 | 57.418 | 8.3 | 80.427 | 10.9 | 95.848 | 13.5 | 95.114 | 16.1 | 86.026 | 18.7 | 67.891 |
| 3.2 | 32.560 | 5.8 | 53.377 | 8.4 | 81.196 | 11.0 | 96.207 | 13.6 | 97.894 | 16.2 | 85.355 | 18.8 | 67.262 |
| 3.3 | 33.573 | 5.9 | 59.334 | 8.5 | 81.956 | 11.1 | 96.547 | $: 3.7$ | 37.654 | 16.3 | 84.677 | 18.9 | 56.643 |
| 3.4 | 34.579 | 5.0 | 69.265 | 8.6 | 82.703 | 11.2 | 96.869 | 13.8 | 97.387 | 16.4 | 83.990 | 19.0 | 65.038 |
| 3.5 | 35.355 | 0.1 | 61.232 | 8.7 | 83.440 | 11.3 | 97.168 | 13.9 | 97.101 | 16.5 | 83.319 | 19.1 | 65.447 |
| 3.6 | 3x. xC5 | 6.2 | 52.176 | B. 3 | 84.164 | 11.4 | 97.445 | 14.0 | 95.793 | 16.6 | 82.594 | 19.2 | 64.871 |
| 3.7 | 37.812 | 6.3 | 63.115 | 8.9 | 84.875 | $: 1.5$ | 97.702 | 14.1 | 76.464 | 16.\% | 81.83S | 19.3 | 64.305 |
| 3.8 | 38.622 | 6.4 | 64.046 | 9.0 | 85.567 | 11.6 | 97.938 | 14.2 | 96.113 | 16.8 | 81.179 | 19.4 | 63.758 |
| 3.9 | 39.629 | 6.5 | 64.975 | 9.1 | 86.250 | 11.7 | 78.149 | 14.3 | 95.740 | 16.9 | 80.451 | 19.5 | 63.222 |
| 4.0 | 40.633 | 5.6 | 65.896 | 9.2 | 86.323 | 11.8 | 98.342 |  | 95.348 | 17.0 | 79.742 | 119.6 | 62.703 |
| 4.1 | 41.637 | 6.7 | 46.810 | 9.3 | 87.5\%9 | 12.9 | 98.510 | 14.5 | 94. 236 | 1:7.1 | 79.023 | . 7 | 52.201 |
| 4.2 | 42.645 | 6.3 | 67.720 | 7.4 | 88.221 | 12.0 | 98.658 | 14.6 | 94.504 | 17.2 | 78.301 | . 8 | 61.714 |
| 4.3 | 43.639 | 6.9 | 60.623 | 9.5 | 83.844 | 12.1 | 98.783 | 14.7 | 94.054 | 17.3 | 77.578 | 19.9 | 61.243 |
| 4.4 | 44.641 | 7.0 | 69.518 | 9.5 | 49.450 | 12.2 | 98.882 | 14.8 | 23.565 | 17.4 | 76.654 | 20.0 | 60,788 |
| 4.5 | 45.639 | 7.1 | 70.707 | 9.7 | 90.053 | 12.3 | 98.965 | 14.9 | 73.095 | 17.5 | 76.134 |  |  |



Fig. A-4. E- and H-plane field patterns


(d) 3.20 cm , H-plane

Fig. A-4. E- and H-plane field patterns

(e) 6.67 cm, E-plane

(f) $6.67 \mathrm{~cm}, \mathrm{H}$-plane

Fig. A-4. E- and H-plane field patterns

TABLE A-2
Summary of Gain-Standard Horn Data

| Band | Frequency Range | $\begin{gathered} \text { Dimensions (I.D.) } \\ \text { (in.) } \end{gathered}$ | DesignPoint Frequency | Gain at Design Point (db) |
| :---: | :---: | :---: | :---: | :---: |
| $\int 8 \mathrm{~mm}$ | $\left.\begin{array}{r} 0.77-1.13 \mathrm{~cm} \\ 26,550-38,960 \mathrm{Mc} \end{array} \right\rvert\,$ | $\begin{aligned} & a=2.720 b=2.231 \\ & \ell_{H}=6.513 \quad \ell_{E}=6.197 \end{aligned}$ | $\begin{array}{r} 0.85 \mathrm{~cm} \\ 35,290 \mathrm{Mc} \end{array}$ | 24.7 |
| $\{1.25 \mathrm{~cm}$ | $\left\|\begin{array}{r} 1.13-1.66 \mathrm{~cm} \\ 18,070-26.550 \mathrm{Mc} \end{array}\right\|$ | $\begin{aligned} & a=4.000 \quad b=3.281 \\ & \ell_{\mathrm{H}}=9.706 \quad \ell_{\mathrm{E}}=9.113 \end{aligned}$ | $\begin{array}{r} 1.25 \mathrm{~cm} \\ 24,000 \mathrm{Mc} \end{array}$ | 24.7 |
| ( 1.8 cm | $\left\|\begin{array}{r} 1.66-2.42 \mathrm{~cm} \\ 12.400-18,070 \mathrm{Mc} \end{array}\right\|$ | $\begin{array}{ll} a=5.984 & b=4.908 \\ \ell_{H}=14.333 & \ell_{E}=13.633 \end{array}$ | $\begin{array}{r} 1.87 \mathrm{~cm} \\ 16,040 \mathrm{Mc} \end{array}$ | 24.7 |
| $\int 3.2 \mathrm{~cm}$ | $\begin{aligned} & 2.42-3.70 \mathrm{~cm} \\ & 8100-12,400 \mathrm{Mc} \end{aligned}$ | $\begin{array}{ll} a=7.654 & b=5.669 \\ \ell_{\mathrm{H}}=13.484 \quad \ell_{\mathrm{E}}=12.598 \end{array}$ | $3.20 \mathrm{~cm}$ $9375 \mathrm{Mc}$ | 22.1 |
| $\{4.75 \mathrm{~cm}$ | $\begin{aligned} & 3.60-5.20 \mathrm{~cm} \\ & 5770-8330 \mathrm{Mc} \end{aligned}$ | $\begin{array}{ll} a=11.360 & b=8.415 \\ \ell_{H}=20.014 & \ell_{E}=18.700 \end{array}$ | 4.75 cm 6315 Mc | 22.1 |
| $\int 3.95 \mathrm{~cm}$ | $\begin{aligned} & 3.00-4.30 \mathrm{~cm} \\ & 6980-10,000 \mathrm{Mc} \end{aligned}$ | $\begin{array}{ll} a=5.041 & b=3.733 \\ \ell_{\mathrm{H}}=7.447 & \ell_{\mathrm{E}}=6.555 \end{array}$ | $\begin{aligned} & 3.95 \mathrm{~cm} \\ & 7595 \mathrm{Mc} \end{aligned}$ | 18.0 |
| $\{6 \mathrm{cmi}$ | $\begin{aligned} & 5.10-7.60 \mathrm{~cm} \\ & 3950-5880 \mathrm{Mc} \end{aligned}$ | $\begin{array}{ll} a=8.507 & b=6.300 \\ l_{H}=12.462 & \ell_{E}=11.062 \end{array}$ | $\begin{aligned} & 6.67 \mathrm{~cm} \\ & 4500 \mathrm{Mc} \end{aligned}$ | 18.0 |
| 10 cm | $\begin{aligned} & 7.60-11.5 \mathrm{~cm} \\ & 2600-\quad 3950 \mathrm{Mc} \end{aligned}$ | $\left\lvert\, \begin{array}{ll} a=12.760 & b=9.450 \\ \ell_{H}=18.682 & \ell_{E}=16.593 \end{array}\right.$ | $\begin{array}{r} 10.00 \mathrm{~cm} \\ 3000 \mathrm{Mc} \end{array}$ | 18.0 |
| $\int 15 \mathrm{~cm}$ | $\begin{aligned} & 11.5-17.6 \mathrm{~cm} \\ & 1700-2600 \mathrm{Mc} \end{aligned}$ | $\left\|\begin{array}{ll} a=14.508 & b=10.747 \\ \ell_{\mathrm{H}}=16.508 & \ell_{\mathrm{E}}=14.107 \end{array}\right\|$ | $\begin{array}{r} 15.22 \mathrm{~cm} \\ 1970 \mathrm{Mc} \end{array}$ | 15.5 |
| 23 cm | $\begin{aligned} & 17.6-\quad 26.5 \mathrm{~cm} \\ & 1130-\quad 1700 \mathrm{mc} \end{aligned}$ | $\left.\begin{array}{ll} a=21.931 & b=16.245 \\ l_{\mathrm{H}}^{\prime}=24.955 & \ell_{\mathrm{E}}=21.325 \end{array} \right\rvert\,$ | $\begin{array}{r} 23.00 \mathrm{~cm} \\ 1300 \mathrm{Mc} \end{array}$ | 15.5 |
| 30 cm | $\begin{array}{r} 26.0-31.5 \mathrm{~cm} \\ 950-1150 \mathrm{Mc} \end{array}$ | $\begin{array}{ll} a=21.931 & b-16.245 \\ \ell_{\mathrm{H}}=28.730 & \ell_{\mathrm{E}}=24.000 \end{array}$ | $\begin{aligned} & 30.00 \mathrm{~cm} \\ & 1000 \mathrm{Mc} \end{aligned}$ | 13.7 |

Horns in brackets are scaled versions of each other, excer for the $\boldsymbol{\ell}_{H}$ dimensions, which are chosen to make a simple butt-joint at the waveguide


Fig. A-5 (a). Gain curves


Fig. A-5 (b). Gain curves


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Fig. A-5 (c). Gain curves and conversion chart

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Fig. A-6. Electroformed horn, 8-mm-band gain-atandard ( 0.77 -1.13 cm)

Fig. A-z Mandril for electroforming 8-mm-band gain-standard horn

Fig. A-8. 1.23-cm-band gain-standard horn (1.13-1.66 cm)

Fig. A-9. $18-\mathrm{mm}-\mathrm{band}_{\mathrm{g}} \mathrm{galn}-\mathrm{standard}$ horn ( $1.66-2.42 \mathrm{~cm}$ )

Fig. A-10. $3.2-\mathrm{cm}$-band gain-stindard horn (2.42-3.70 cm)

Fig. A. $11 . \quad 3.95-\mathrm{cm}$-band gain-standard horn ( 3.0 .4 .30 cm )

Fig. A-12. 4.75-cm-band gain-standard horn (3.60-5.20 cmi)

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Fig. A-14. 10 -cm-band gain-standard horn (7.60.11.5 cm )


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15-cm-band gair-standard horn (11.5-17.6cm)


Fig. A-16. 23-cm-band gain-standard horn ( $17.6-26.5 \mathrm{~cm}$ )

Fig. A-17. 30 -cm-band gain-standard horn ( $26,0.31 .5 \mathrm{~cm}$ )

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[^0]:    *With the exception of Fig. A-1, all figures and tables bearing the letter $A$ nre grouped at tite end af the A.ppendix, and are listed on page 6.
    $\dagger_{F o r}$ a general description of the methods used in making such measurements see Footmo +2 , p. 7 of the Appendix, raf. pp. 582-58f. The remarks in this referonce about the minimum meparation digiance fos the ho: is should be re-evalueted in the llyitt of Ref. 2.

[^1]:    ${ }^{1}$ Schelkunoff, S. A., "Electromagnetic Waves," D. Van Nostrand, Inc.. New Mork, po. 363-365. 1943
    Zsilver, S.: "Microwave Antena Thecry \& Design,"MrGravr-Hill Book Co., Iric., New York, pp. 595-589, 1949
    ${ }^{3}$ Schelkanoff, S. A., and Friln, H. T., "Antennas. Theory arid Pracilce," John Wiley and Sons, Inc.. New York, pp. 528-529, 1952

[^2]:    4Braun, E. H. "Calculation of the Gain of Small Horns," Proc. I,R,E., Vol. 41. No. 12. pp. 178S-6, Lec. 1953

[^3]:    *An optirrum horn is one for which the aperture dimensions have been chosen to give
     and $\mathrm{t}^{2} \cong 2.08 \lambda \mathrm{E}_{\mathrm{E}}$
    This has becn worked out by E. H. Braun in an uripublished raport.

